Superconductivity and Mottness: Exact Results

Nature Physics, vol.16, 1175-1180 (2020) with N&V by J. Zaanen

arxiv.org/abs/2103.03256

Luke Yeo



Edwin Huang

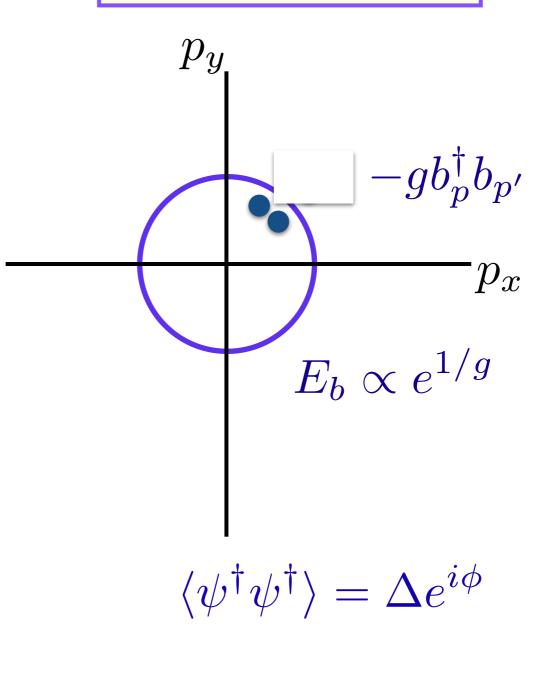
G. La Nave

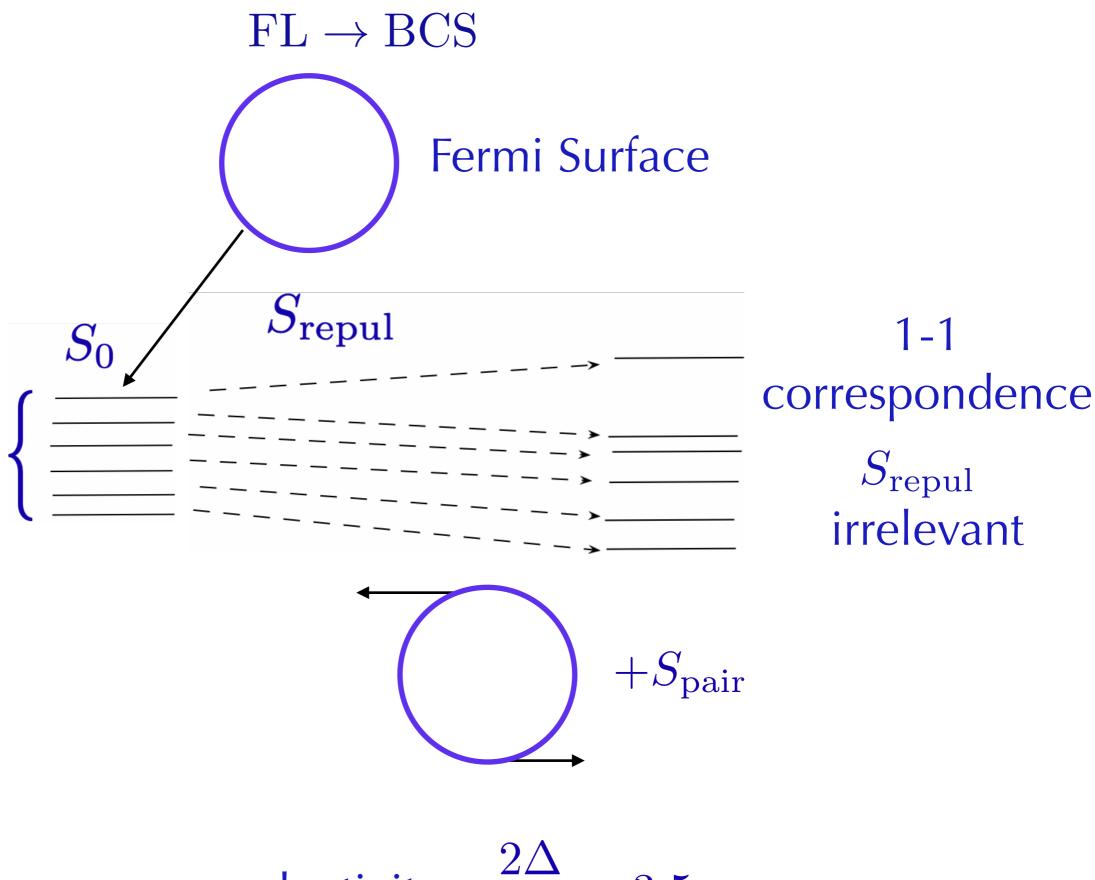




Kammerlingh Onnes (Leiden) · LIQUIFIES HELIUM 1908 DISCOVERS SUPERCONDUCTIVING in MERCURY 1911 4125 RESISTANCE (D) Receives NOBEL PRIZE 1913 Hg 9075 0,05 0,025 Figure 1 Resistance in ohms of a specimen of mercury versus absolute temperature. 10-5 M This plot by Kamerlingh Onnes marked the discovery of superconductivity. 4.0 TEMPERATURE ("K)

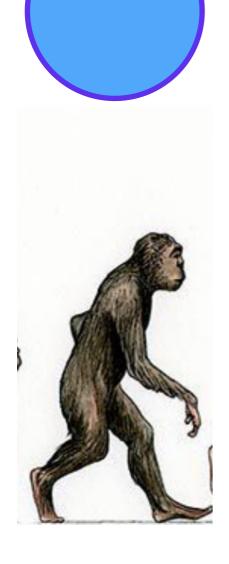
Cooper instability





superconductivity
$$\frac{2\Delta}{T_c} = 3.5$$

Fermi gas



Fermi liquid



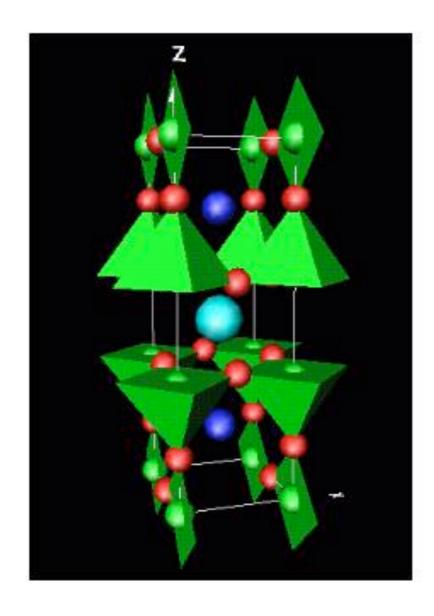
BCS superconductor



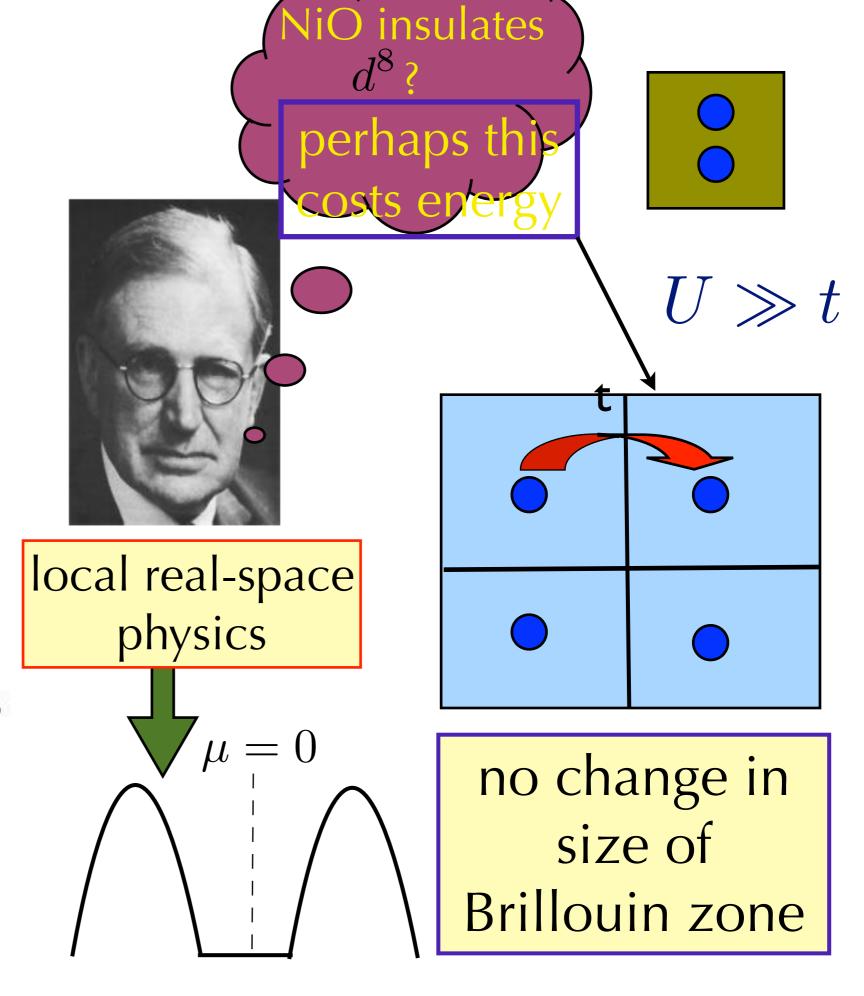


Is there physics beyond BCS?

fixed point beyond FL?



Y Ba₂Cu₃O₇ Cuprate Superconductors



solve the Hubbard Model!!

Cooper instability??

Progress thus far?

DMFT

QMC

disputes

Sept. 1997

Nov. 1997

A Critique of "A Critique of Two Metals" A Critique of Two Metals

R. B. Laughlin

Departrment of Physics

Stanford University

Stanford, California 94305

idea is either missing or improperly understood. Another indicator that something is deeply wrong is the inability of anyone to describe the elementary excitation spectrum of the Mott insulator precisely even as pure phenomenology. Nowhere can one find a quantitative band structure of the elementary particle whose spectrum becomes gapped. Nowhere can one find precise information about the particle whose gapless spectrum causes the paramagnetism. Nowhere can one find information about the interactions among these particles or of their potential bound state spectroscopies. Nowhere can one find precise definitions of Mott insulator terminology. The upper and lower Hubbard bands, for example, are vague analogues of the valence and conduction bands of a semiconductor, except that they coexist and mix with soft magnetic excitations no one knows how to describe very well.

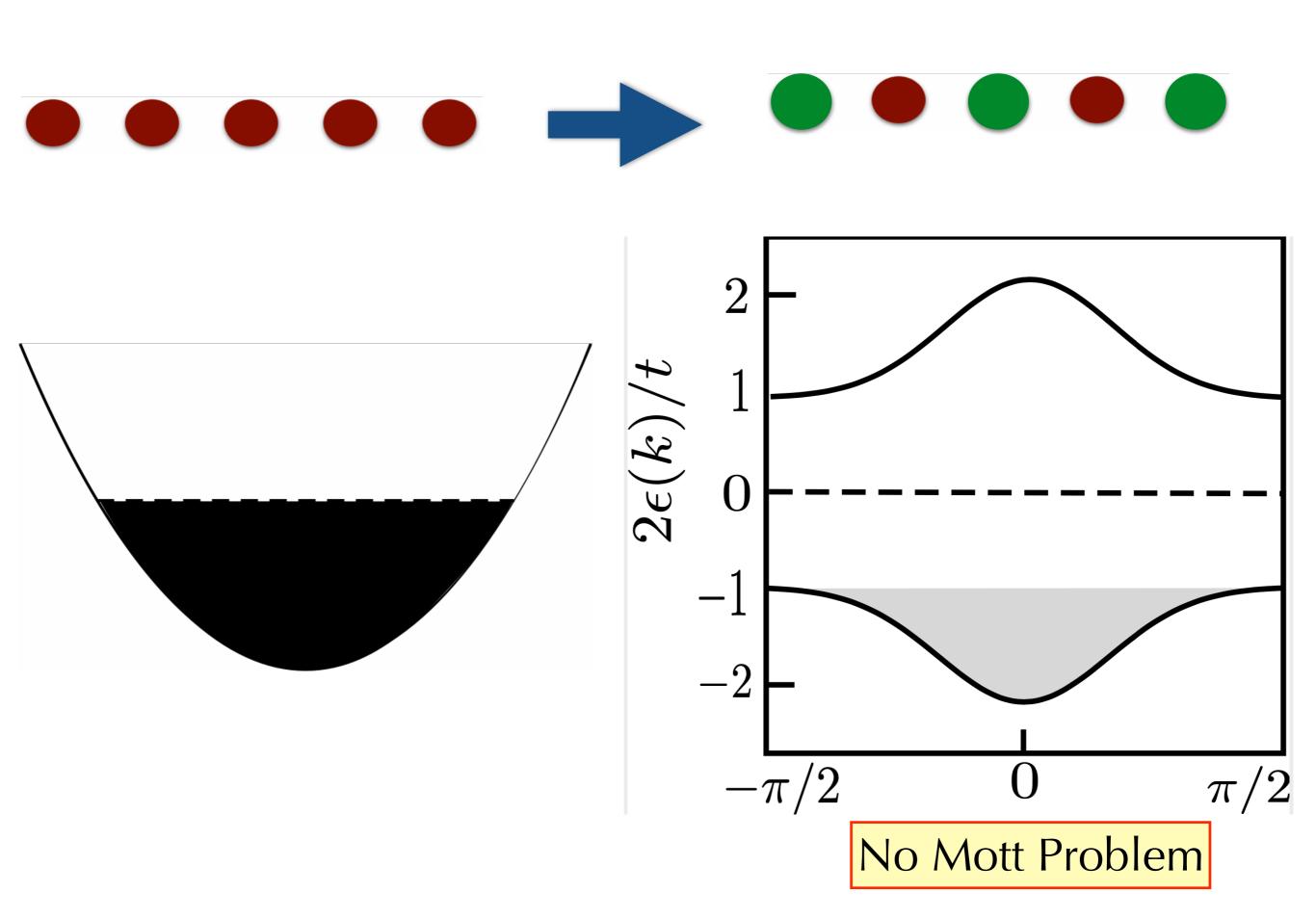
Philip W. Anderson and G. Baskaran

Joseph Henry Laboratories of Physics

Princeton University, Princeton, NJ 08544

The fundamental argument is presented in the second paragraph: "Ten years of work by some of the best minds in theoretical physics have failed to produce any formal demonstration"... of the Mott insulating state. The statement would be ludicrous if it were not so influential. The proviso "at zero temperature" is added, because of course most Mott

concern. It is the tragedy of Mott that although he almost certainly won his Nobel prize for the Mott insulator, Slater, who couldn't think clearly about finite temperature, won the publicity battle.



gap with no symmetry Laughlin's objection: breaking not demonstrated!! μ charge gap 3 $1 \rightarrow 2$





Mottness

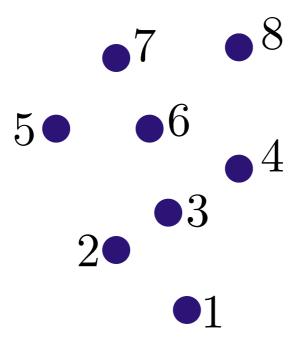
zeros

$$= 0$$

$$\neq 0$$

$$DetReG(\omega = 0, p)$$

counting particles



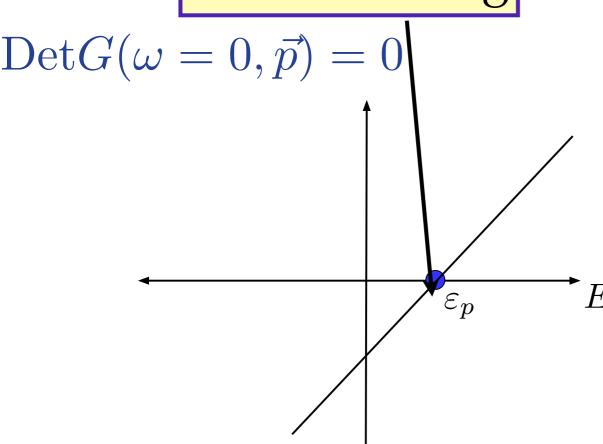
is there a more efficient way?

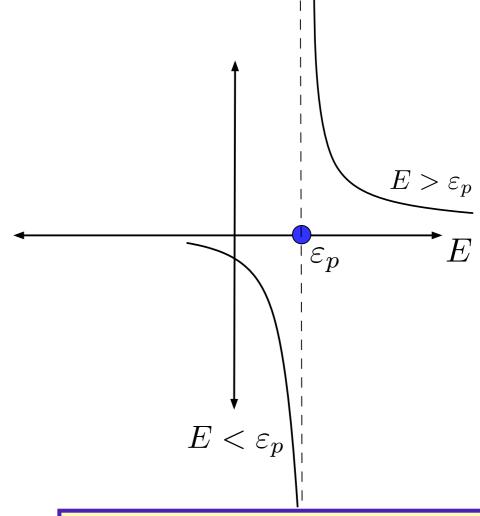
Luttinger counting theorem

$$G(E) = \frac{1}{E - \varepsilon_p}$$

$$n = 2\sum_{\mathbf{k}} \Theta(\Re G(\mathbf{k}, \omega = \mathbf{0}))$$

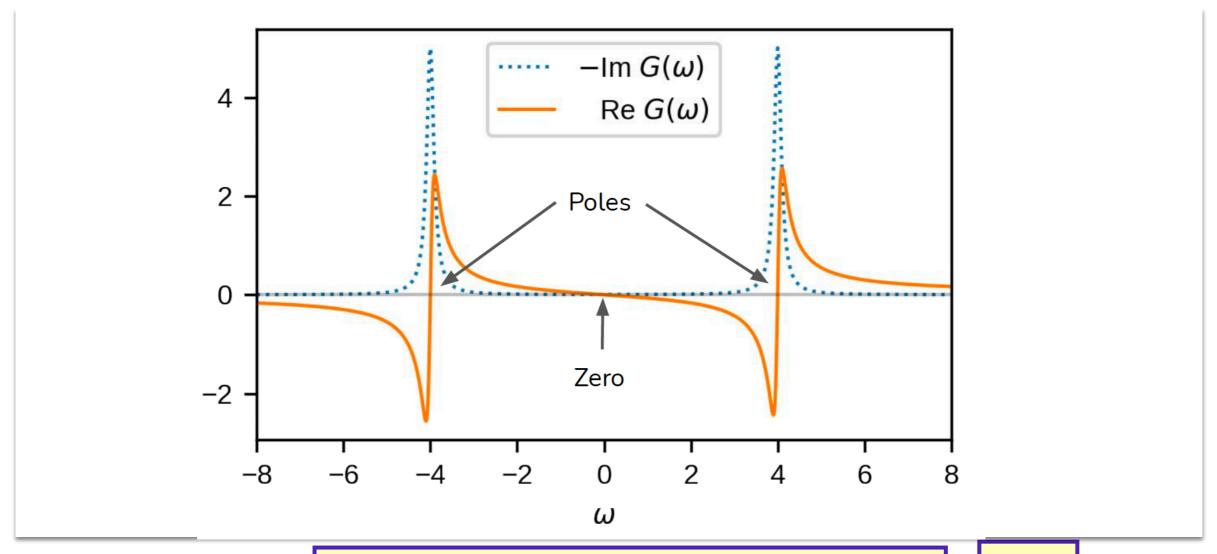
zero-crossing





counting poles (qp)

How do zeros obtain?

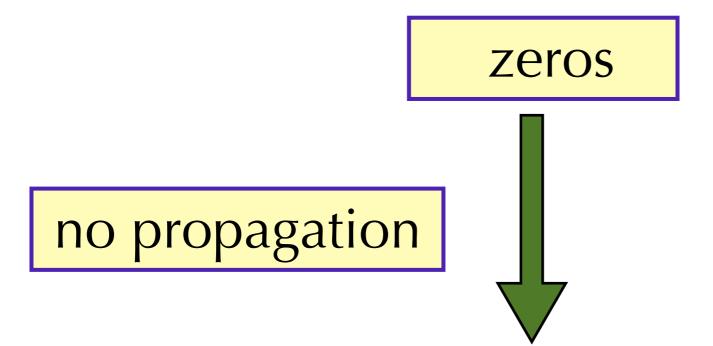


below gap+above gap

= 0

 $\operatorname{DetReG}(k, \omega = 0) = 0$ (single band)

strongly correlated gapped systems



breakdown of particle concept

Mottness

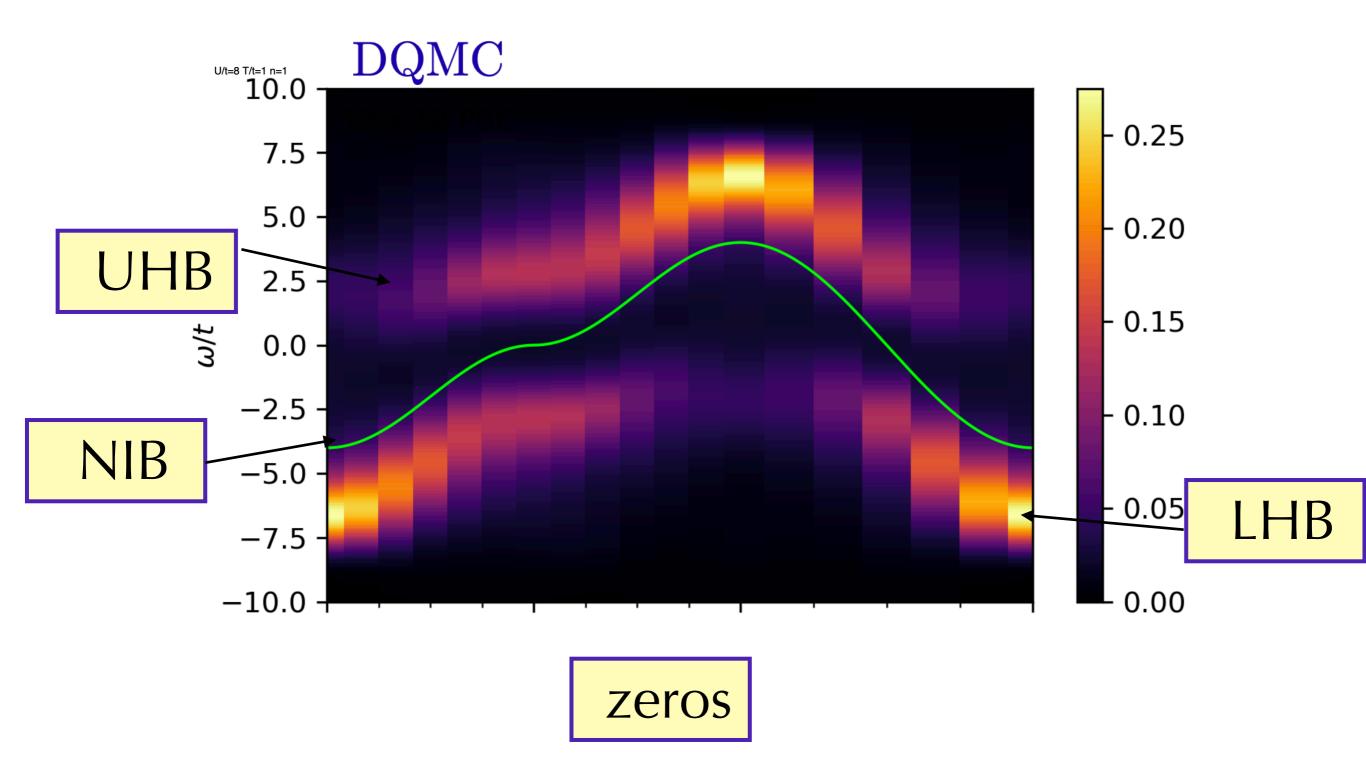
Symmetry Breaking

$$G_k(\omega) = \begin{pmatrix} \frac{1}{\omega - E_k^+} & 0 \\ 0 & \frac{1}{\omega - E_k^-} \end{pmatrix} \begin{pmatrix} \frac{2}{\omega} & 1 \\ \frac{2}{\omega} & 0 \\ -1 & -2 \\ -\pi/2 & 0 & \pi/2 \end{pmatrix}$$

$$\operatorname{Det} G \neq 0$$
 no Mottness

Laughlin

Is the Hubbard model necessary?



No!

Minimal model for Mottness?



Anderson Haldane 2000

2 citations

Fermi liquids

$$H = \sum_{p,\sigma} (\epsilon(p) - \epsilon_F) n_{p\sigma} + \cdots$$

 $(n_{p\uparrow}, n_{p\downarrow})$ conserved currents

 $(c_{p\uparrow}, c_{p\downarrow}, \text{h.c.})$ 4 objects

$$Det M = 1$$

$$Det M = -1$$

SO(4) proper rotations

improper rotations 19

$$Det M = \pm 1 \implies Z_2 = O(4) \div SO(4)$$

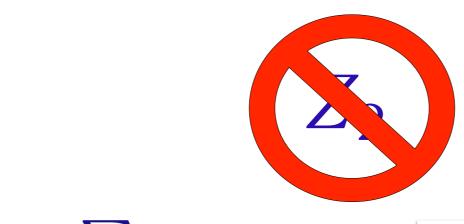
$$Fermi \\ Surface$$

$$H = 0$$

$$n_{p\uparrow} \rightarrow -n_{p\uparrow} \\ n_{p\downarrow} \rightarrow n_{p\downarrow}$$

$$Z_2 \text{ at Fermi } \\ surface \text{ only}$$

How to destroy Fermi liquids?



$$H = \sum_{p,\sigma} \left(\epsilon(p) - \epsilon_F \right) n_{p\sigma} + U n_{p\uparrow} n_{p\downarrow}$$

odd under \mathbb{Z}_2

scaling dimension

$$[n_{p\uparrow}n_{p\downarrow}] = -2$$

relevant interaction

New fixed point!

Hatsugai-Kohmoto model Hubbard not necessary!

Hatsugai-Kohmoto Model (1992)

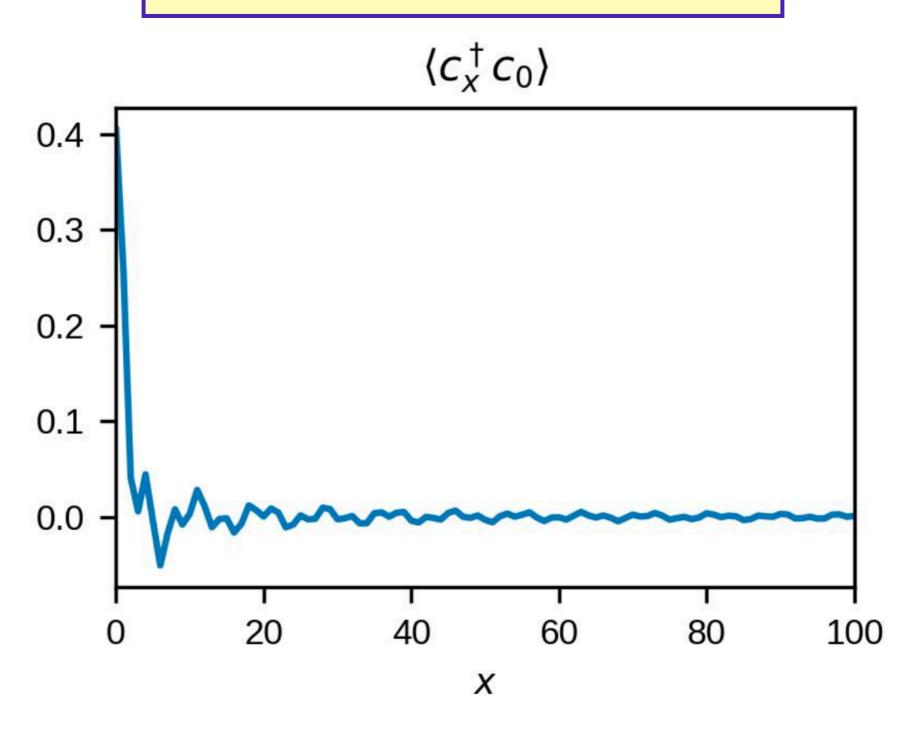
$$H_{\rm HK} = -t \sum_{\langle j,l\rangle,\sigma} \left(c_{j\sigma}^{\dagger} c_{l\sigma} + h.c. \right) - \mu \sum_{j\sigma} c_{j\sigma}^{\dagger} c_{j\sigma}$$

$$c_{k\sigma} = \sum_{j} e^{ikj} c_{j\sigma}$$

$$H_{HK} = \sum_{k} H_{k} = \sum_{k} (\xi_{k}(n_{k\uparrow} + n_{k\downarrow}) + Un_{k\uparrow} n_{k\downarrow}).$$

$$\xi_k = \epsilon_k - \mu$$

non-local but correlations are local



General HK Model

$$\sum_{k} (\xi_{k}(n_{k\uparrow} + n_{k\downarrow}) + Un_{k\uparrow} n_{k'\downarrow}).$$
 perturbation

relevant

$$[H_t, H_U] = 0$$

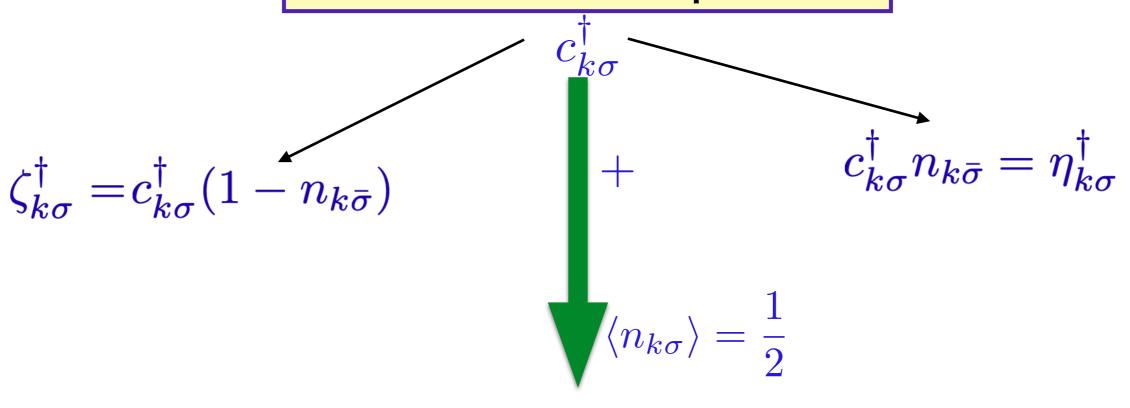
Solvable Mott transition

$$G_{k\sigma}(i\omega_n \to z) = \frac{1 - \langle n_{k\bar{\sigma}} \rangle}{z - \xi_k} + \frac{\langle n_{k\bar{\sigma}} \rangle}{z - (\xi_k + U)}$$

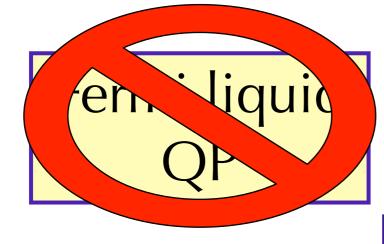
lower Hubbard band

upper Hubbard band

Hubbard band operators



$$G_{\sigma}^{R}(k,\omega) = \frac{1}{\omega + i0^{+} - (\xi_{k} + U/2) - \frac{(U/2)^{2}}{\omega + i0^{+} - (\xi_{k} + U/2)}}$$



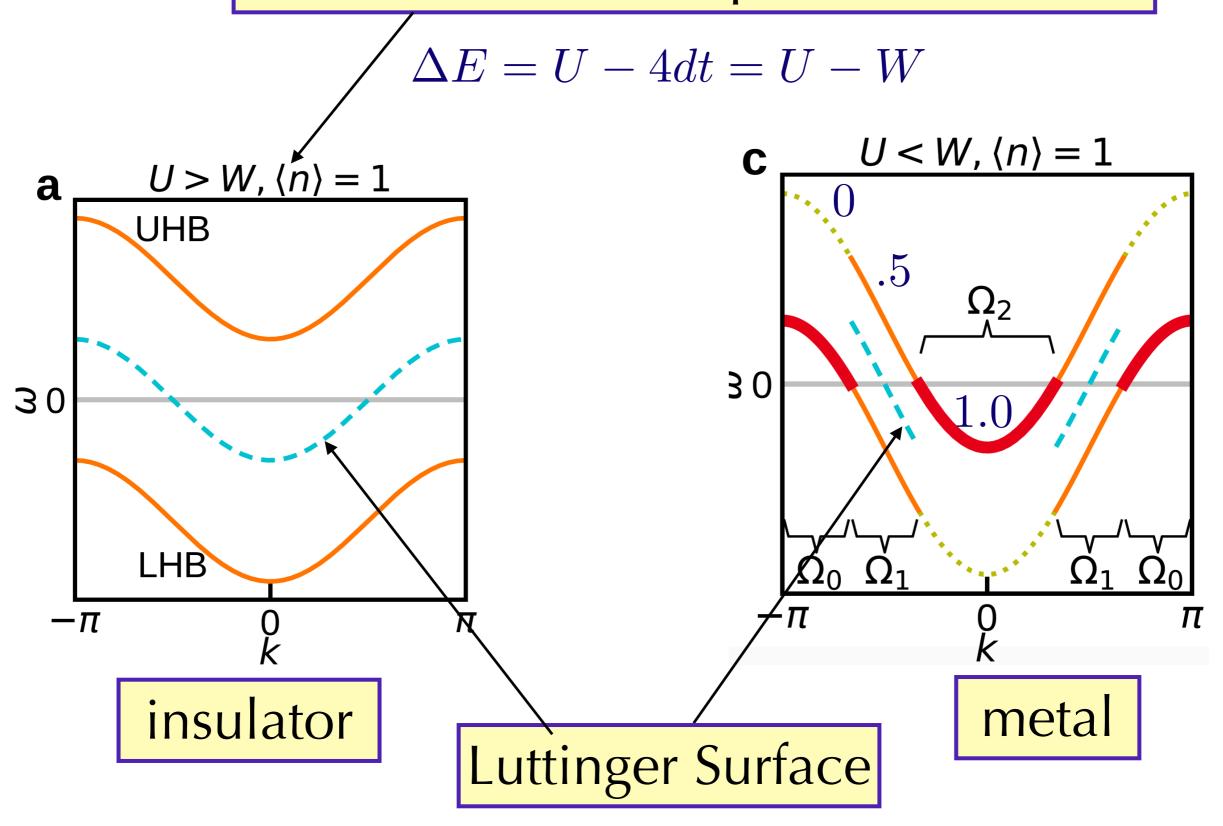
$$\Sigma(k,\omega)$$

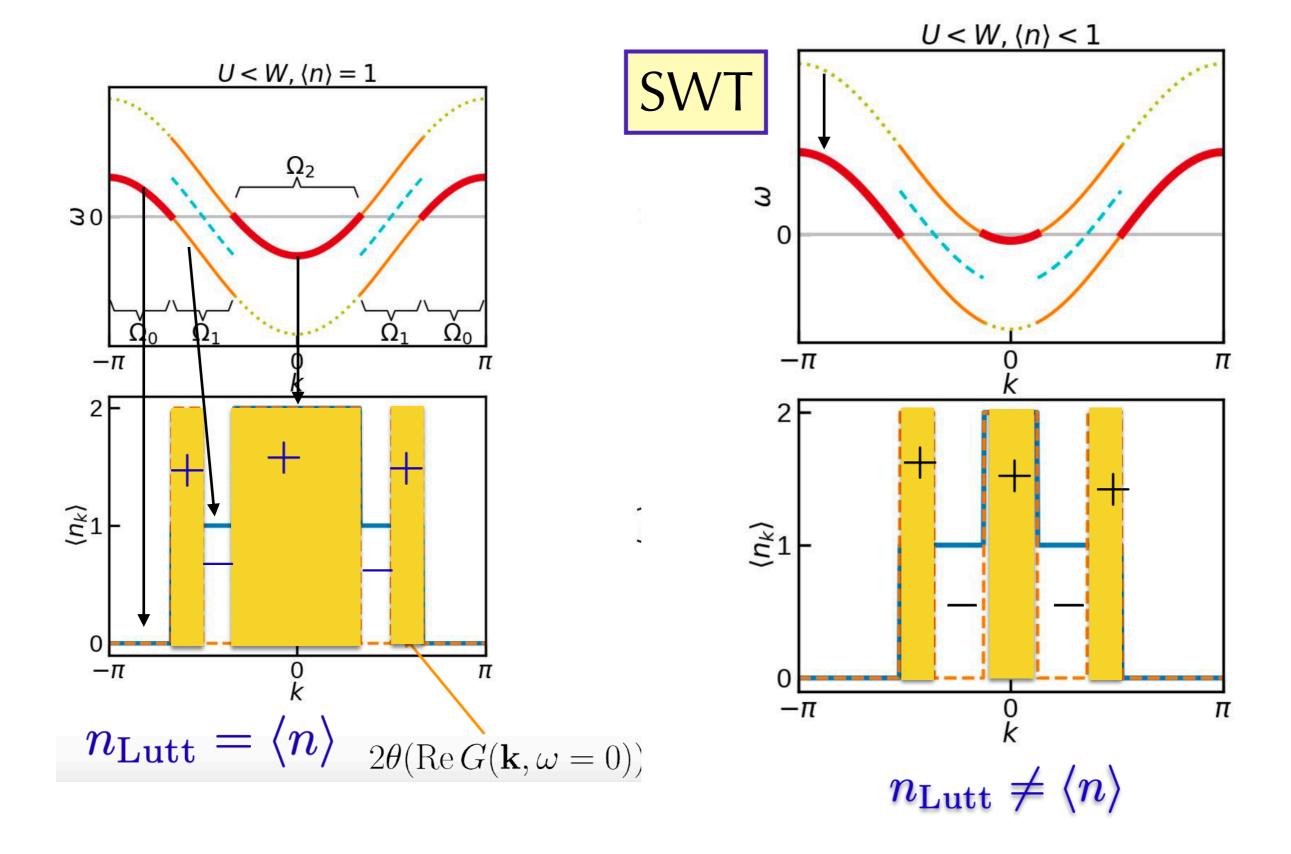
$$\omega = \xi_k + U/2$$

zeros

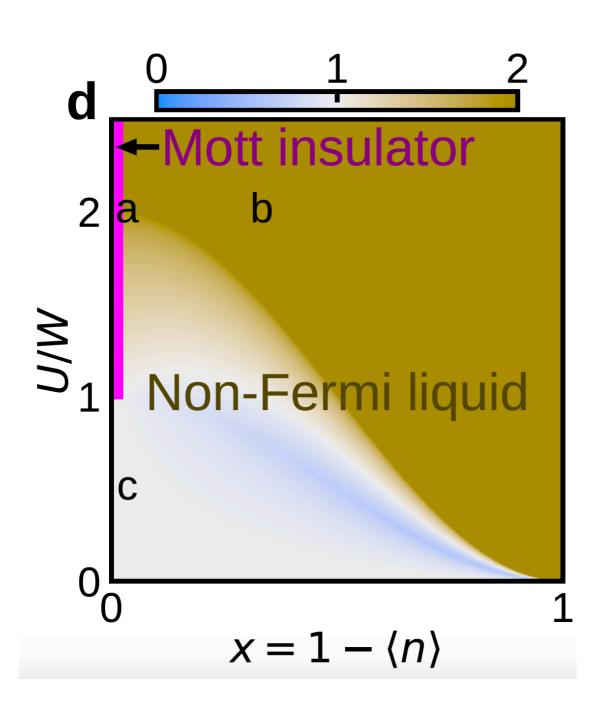
$$\Re \Sigma = \Im \Sigma = \infty$$

Mott transition: composite excitations





Why NFL?



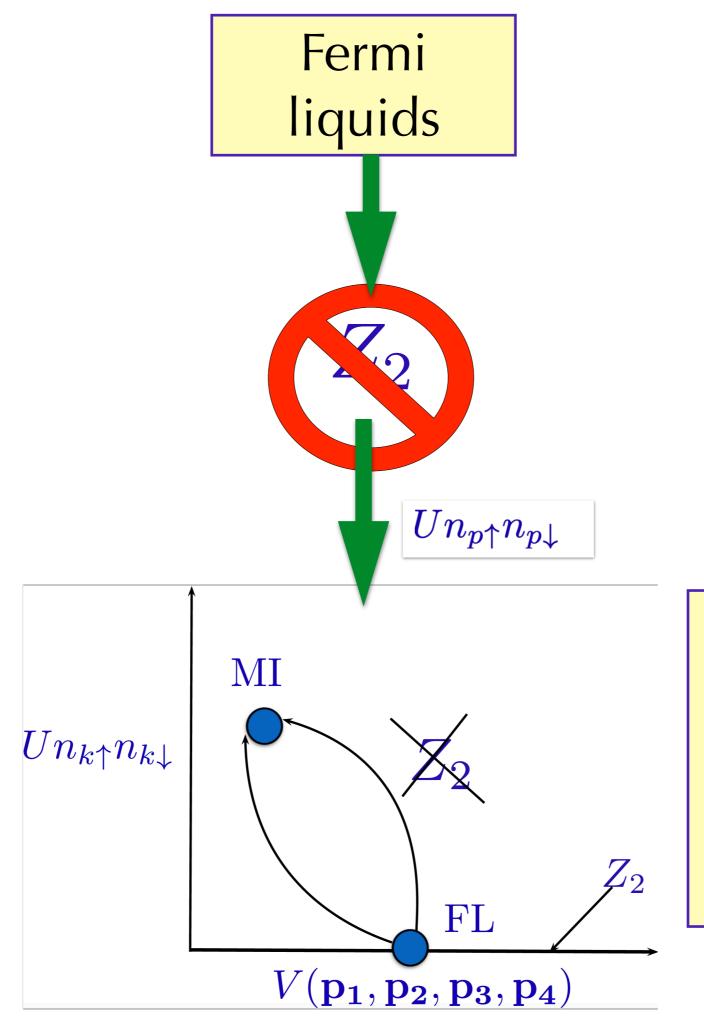
FL

HK Mottness

$$\begin{array}{c} c_{k\uparrow}^{\dagger}c_{k\downarrow}^{\dagger}|G\rangle & ? \\ c_{k\uparrow}^{\dagger}|G\rangle, c_{k\downarrow}^{\dagger}|G\rangle & \zeta_{k\uparrow}^{\dagger}|G\rangle, \zeta_{k\downarrow}^{\dagger}|G\rangle \\ |G\rangle & |G\rangle \end{array}$$

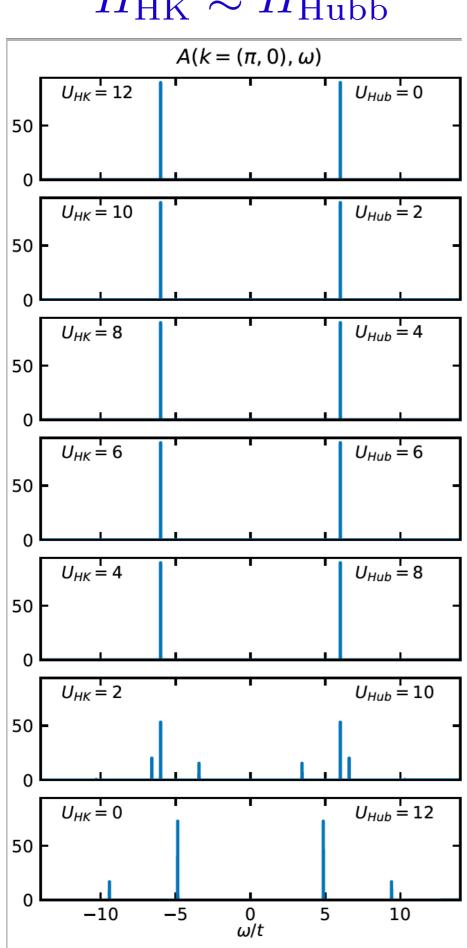
$$\zeta_{k\uparrow}^{\dagger}\zeta_{k\downarrow}^{\dagger}|G\rangle=0$$

EFL?

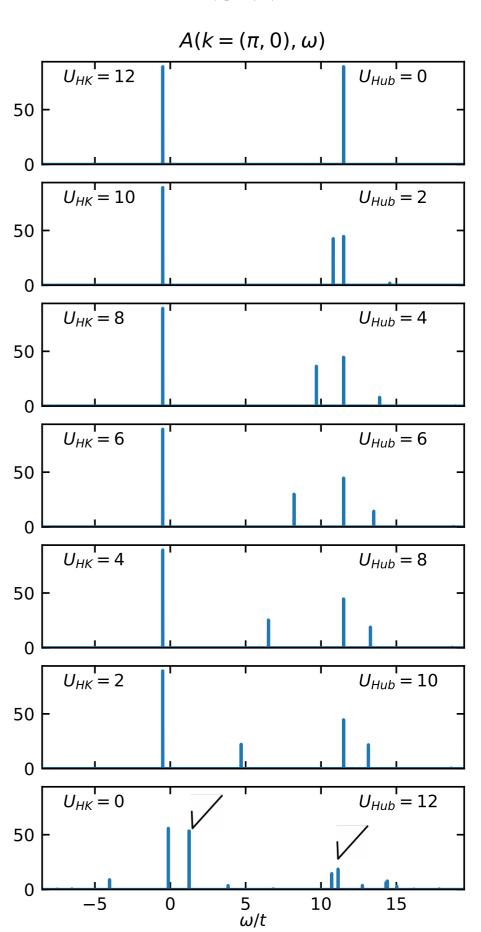


Hubbard not necessary (universality class)

n = 1.0 $H_{\rm HK} \approx H_{\rm Hubb}$

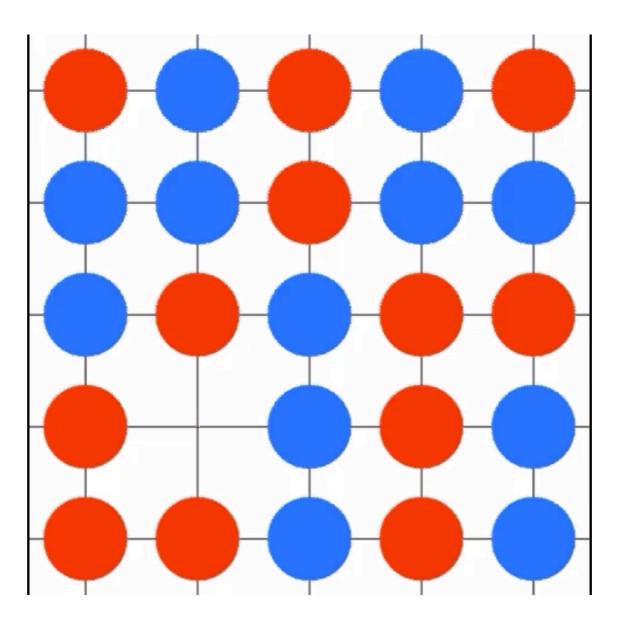


n = 0.875
DSWT



what does the HK model leave out??

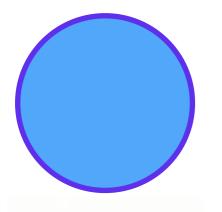
$$[H_t, H_U] \neq 0$$

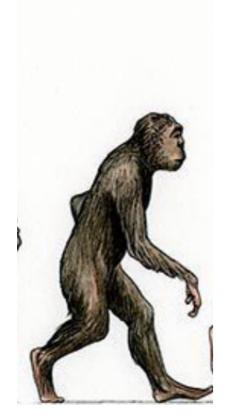


dynamical spectral weight transfer

Mottness

Fermi gas



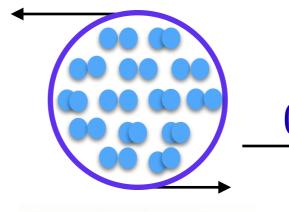


Fermi liquid

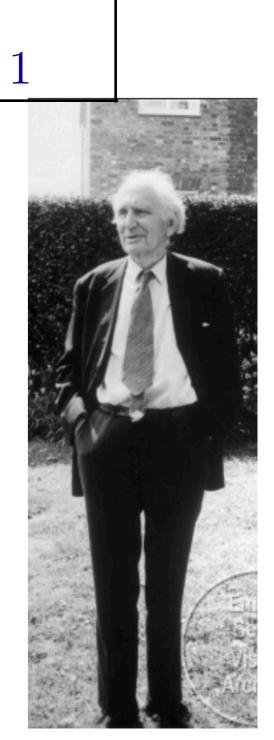




BCS superconductor







Superconductivity?

Cooper Instability

$$H = H_{HK} - gH_{p}$$

$$|\psi\rangle = \sum_{k \in \Omega_{0}} \alpha_{k} b_{k}^{\dagger} |GS\rangle$$

$$k \in \Omega_{0} \qquad \langle n_{k\sigma} \rangle = 0$$

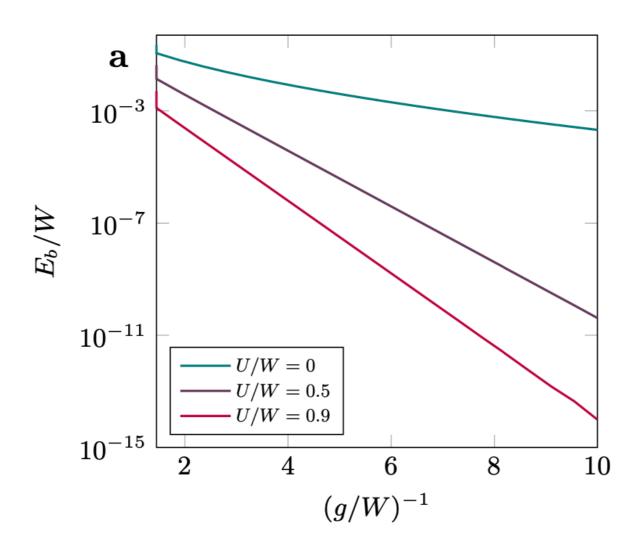
$$E_{b} = \langle GS|H|GS\rangle - \langle \psi|H|\psi\rangle \leq 0$$

$$1 = -\frac{g}{L^d} \sum_{k \in \Omega_0} \frac{\langle 1 - n_{k\uparrow} - n_{-k\downarrow} \rangle}{E - 2\xi_k - U\langle n_{k\downarrow} + n_{-k\uparrow} \rangle}$$

$$1 = -g \int_{\mu}^{W/2} d\epsilon \frac{\rho(\epsilon)}{E - 2\epsilon + 2\mu},$$

Cooper Instability

$$E_b = -E \sim W(1 - (U/W)^2)e^{-\pi W}\sqrt{1 - (U/W)^2}/g$$



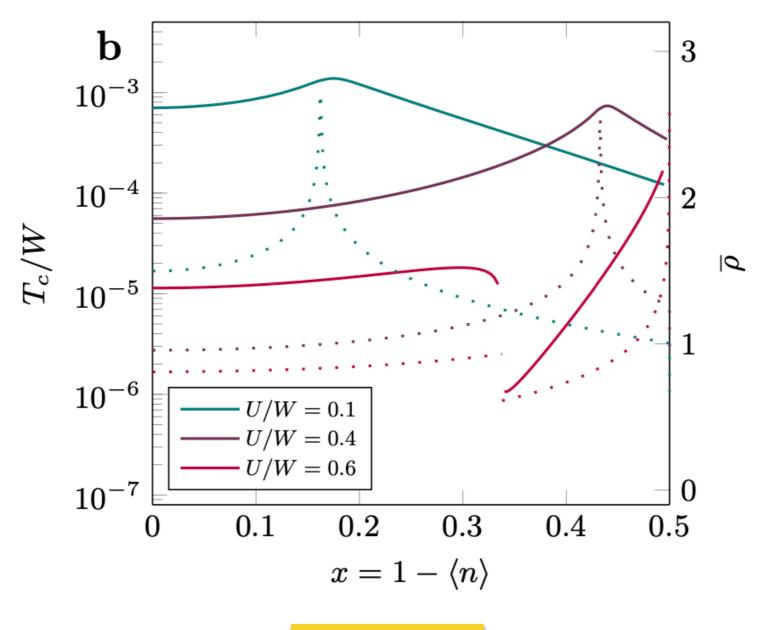
Pair Susceptibility

$$\chi(i\nu_n) \equiv \frac{1}{L^d} \int_0^\beta d\tau e^{i\nu_n \tau} \langle T\Delta(\tau)\Delta^\dagger \rangle_g$$

$$= \frac{\chi_0}{1 - g\chi_0}$$

$$g\chi_0 = 1$$

solve for T_c



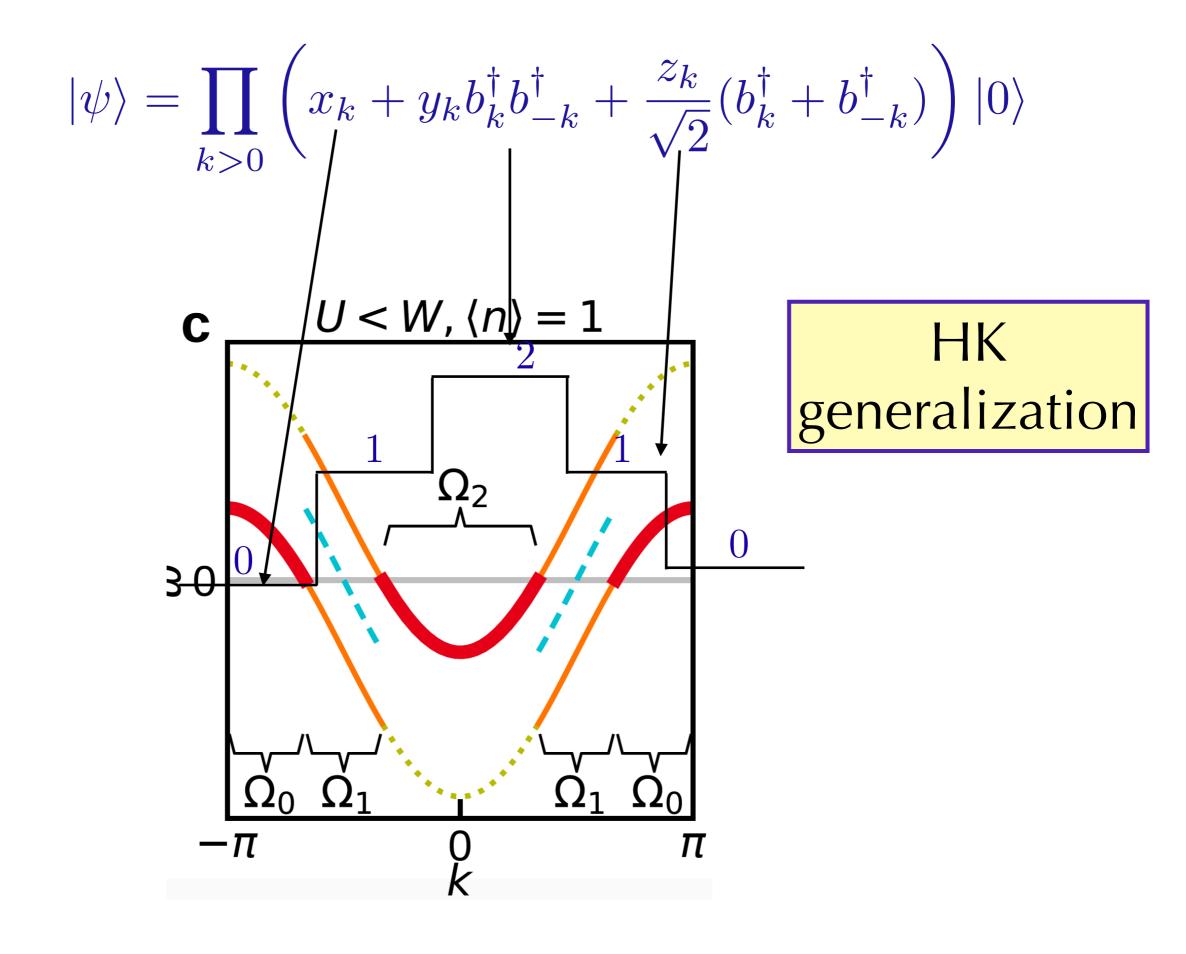
$$W > U$$

$$T_c = (W - U)^{4/5} U^{1/5} \frac{e^{\gamma}}{\pi} e^{-\frac{4}{5} \frac{W}{g}}.$$

variational wave function

$$|\psi_{\text{BCS}}\rangle = \prod_{k} (u_k + v_k b_k^{\dagger} | 0 \rangle$$

$$|\psi_{\text{BCS}}\rangle = \prod_{k>0} (u_k^2 + v_k^2 b_k^{\dagger} b_{-k}^{\dagger} + u_k v_k (b_k^{\dagger} + b_{-k}^{\dagger}))|0\rangle$$

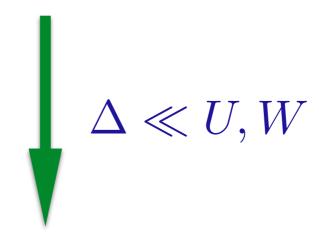


three variational parameters

$$|x_k|^2 + |y_k|^2 + |z_k|^2 = 1$$

gap equation

$$1 = \frac{g}{W} \sinh^{-1}(\frac{W - U}{2\Delta}) + \frac{g}{W} \sinh^{-1}(\frac{U}{2\Delta})$$



$$\Delta = (W - U)^{1/2} U^{1/2} e^{-\frac{W}{2g}}$$

gap/T_c ratio

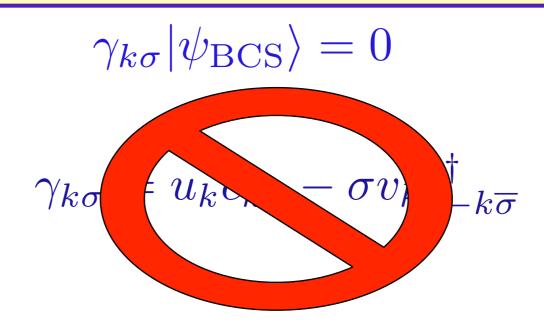
$$\Delta = (W - U)^{1/2} U^{1/2} e^{-\frac{W}{2g}}$$

$$T_c = (W - U)^{4/5} U^{1/5} \frac{e^{\gamma}}{\pi} e^{-\frac{4}{5} \frac{W}{g}}.$$

$$\lim_{g \to 0} \frac{\Delta}{T_c} \to \infty$$

non-BCS superconductivity

Bogoliubov excitations



PYHons excitations

$$\gamma_{k\sigma}^l \propto \sqrt{2}x_k \zeta_{k\sigma}^{\dagger} - \sigma z_k \zeta_{-k\overline{\sigma}}$$

$$\gamma_{k\sigma}^{u} \propto z_{k} \eta_{k\sigma}^{\dagger} - \sigma \sqrt{2} y_{k} \eta_{-k\overline{\sigma}}$$

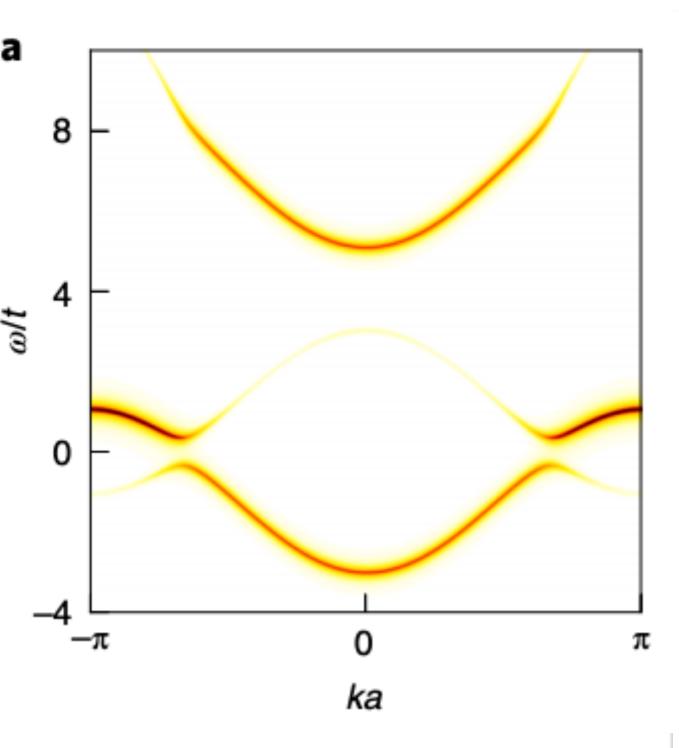
Excitation spectrum

$$\gamma_{k\sigma}^{u/l}|\psi\rangle = 0$$

$$\langle \psi | \gamma_{k\sigma}^{u/l} H \gamma_{k\sigma}^{u/l} \rangle^{\dagger} | \psi \rangle = \langle \psi | H | \psi \rangle + E_k^{u/l}$$

$$E_k^{u/l} = \sqrt{\xi_k^{u/l^2} + \Delta^2}$$

superconductivity affects both bands!



can we explain the color change?

REPORT

Superconductivity-Induced Transfer of In-Plane Spectral

Weight in Bi₂Sr₂CaCu₂O_{8+δ}

H. J. A. Molegraaf¹, C. Presura¹, D. van der Marel^{1,*}, P. H. Kes², M. Li²

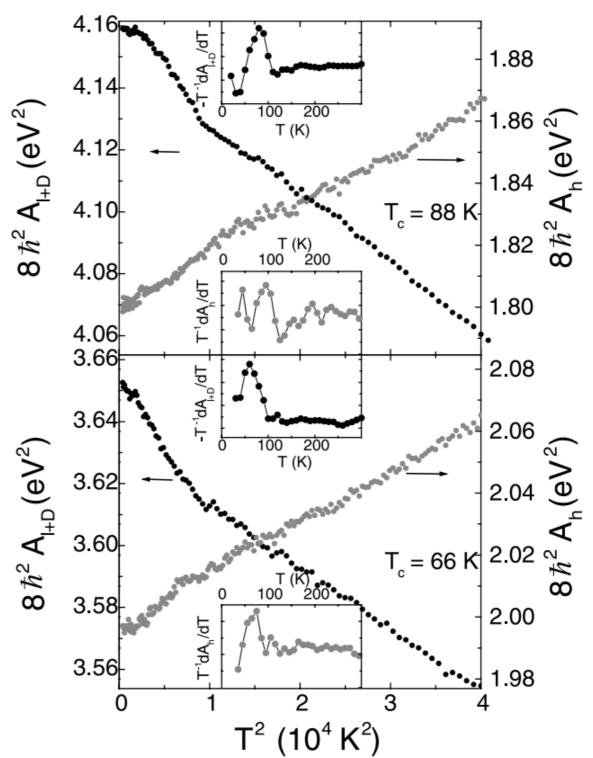
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Science 22 Mar 2002: Vol. 295, Issue 5563, pp. 2239-2241 DOI: 10.1126/science.1069947

$$A_l = \int_0^{\Omega} \sigma(\omega) d\omega \quad \Omega/2\pi c = 10000 cm^{-1}$$

$$A_h = \int_{\Omega}^{2\Omega} \sigma(\omega) d\omega \quad \Omega/2\pi c = 10000 cm^{-1}$$

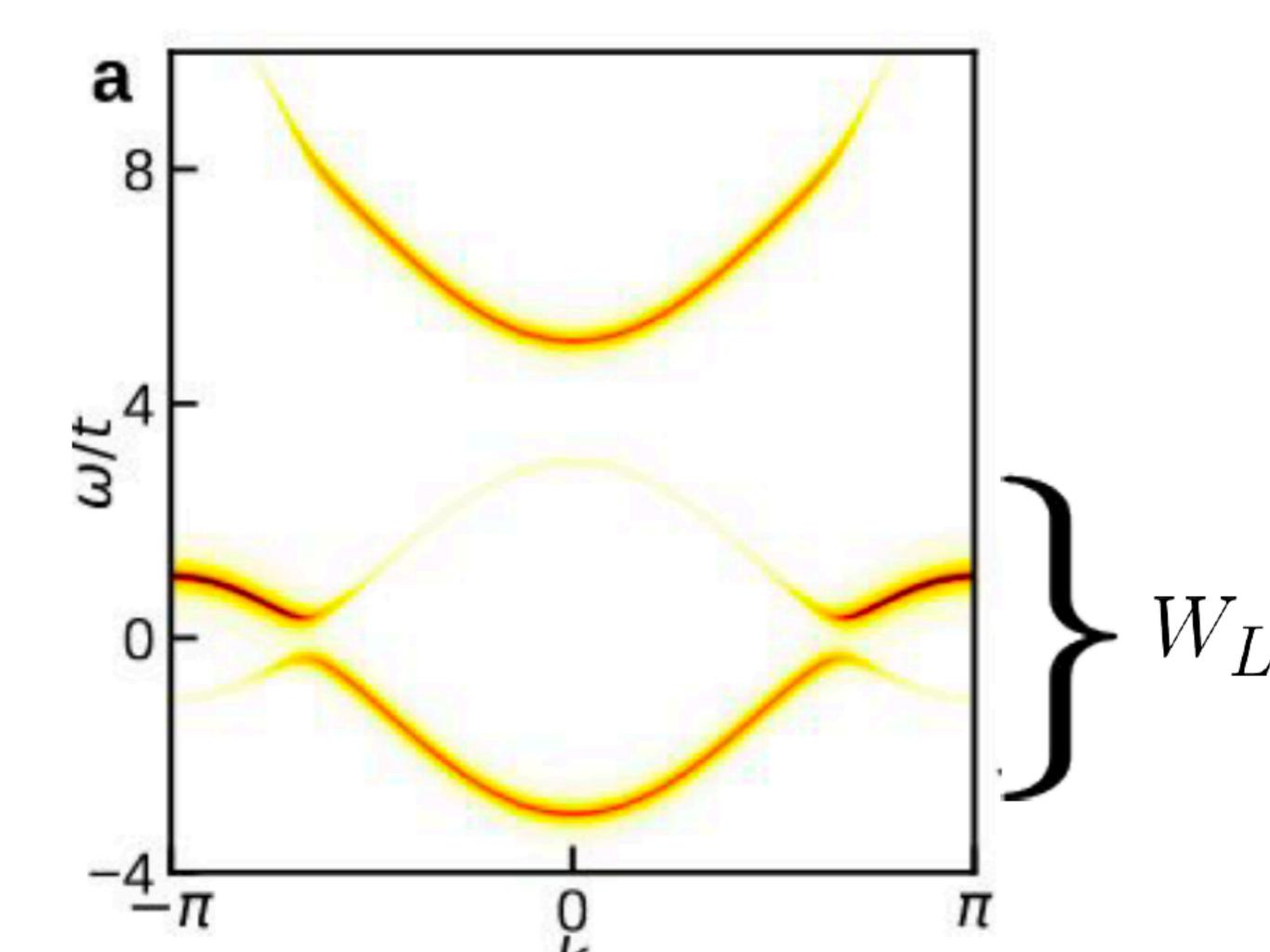
$$rac{\Delta A_l}{A_l} \propto 3\%$$



condensation energy

Optical data are reported on a spectral weight transfer over a broad frequency range of ${\rm Bi_2Sr_2CaCu_2O_{8+8}}$, when this material became superconducting. Using spectroscopic ellipsometry, we observed the removal of a small amount of spectral weight in a broad frequency band from $10^4~{\rm cm^{-1}}$ to at least $2\times10^4~{\rm cm^{-1}}$, due to the onset of superconductivity. We observed a blue shift of the ab-plane plasma frequency when the material became superconducting, indicating that the spectral weight was transferred to the infrared range. Our observations are in agreement with models in which superconductivity is accompanied by an increased charge carrier spectral weight. The measured spectral weight transfer is large enough to account for the condensation energy in these compounds.

UV-IR mixing



why?

$$H = H_{\rm HK} + H_p$$

$$[H_{\rm HK}, H_p] \neq 0$$

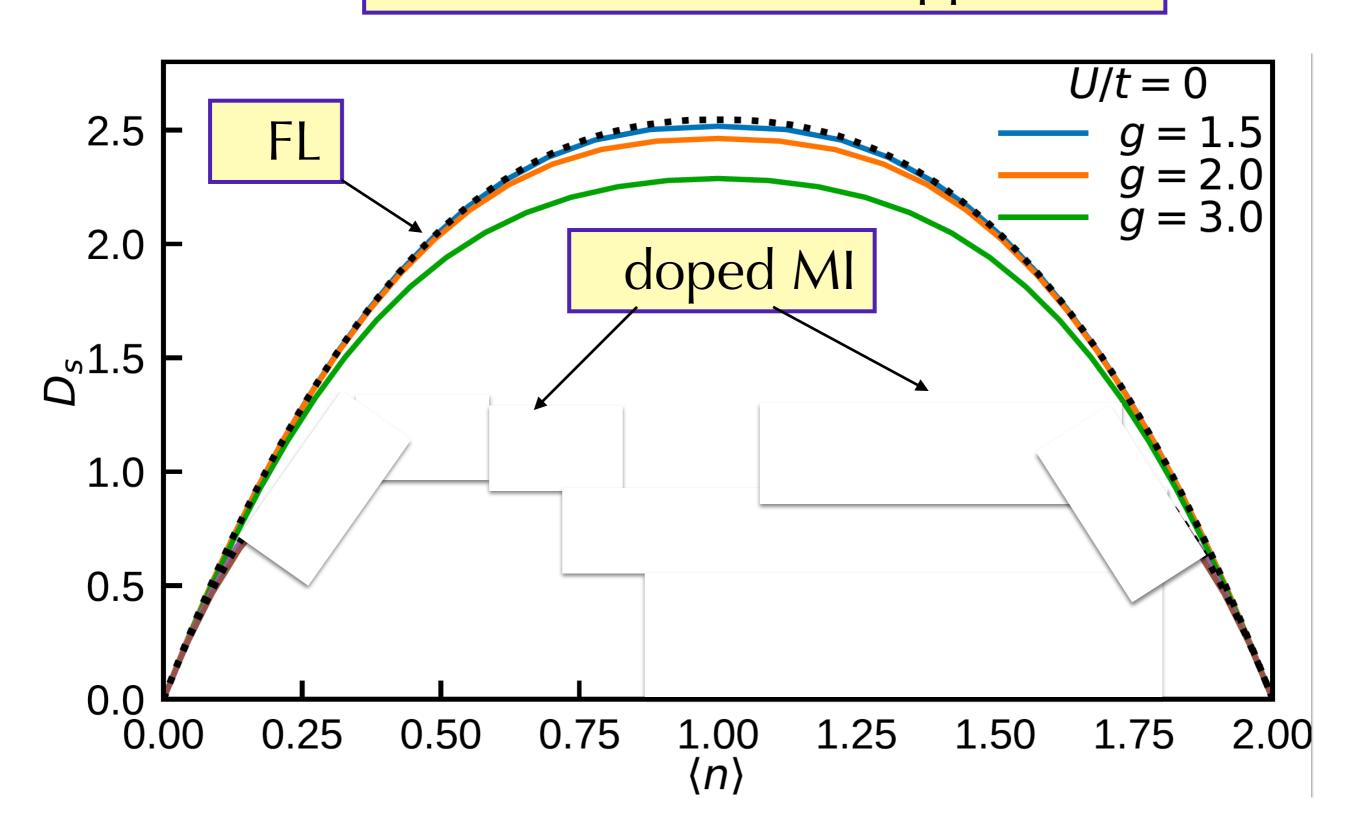


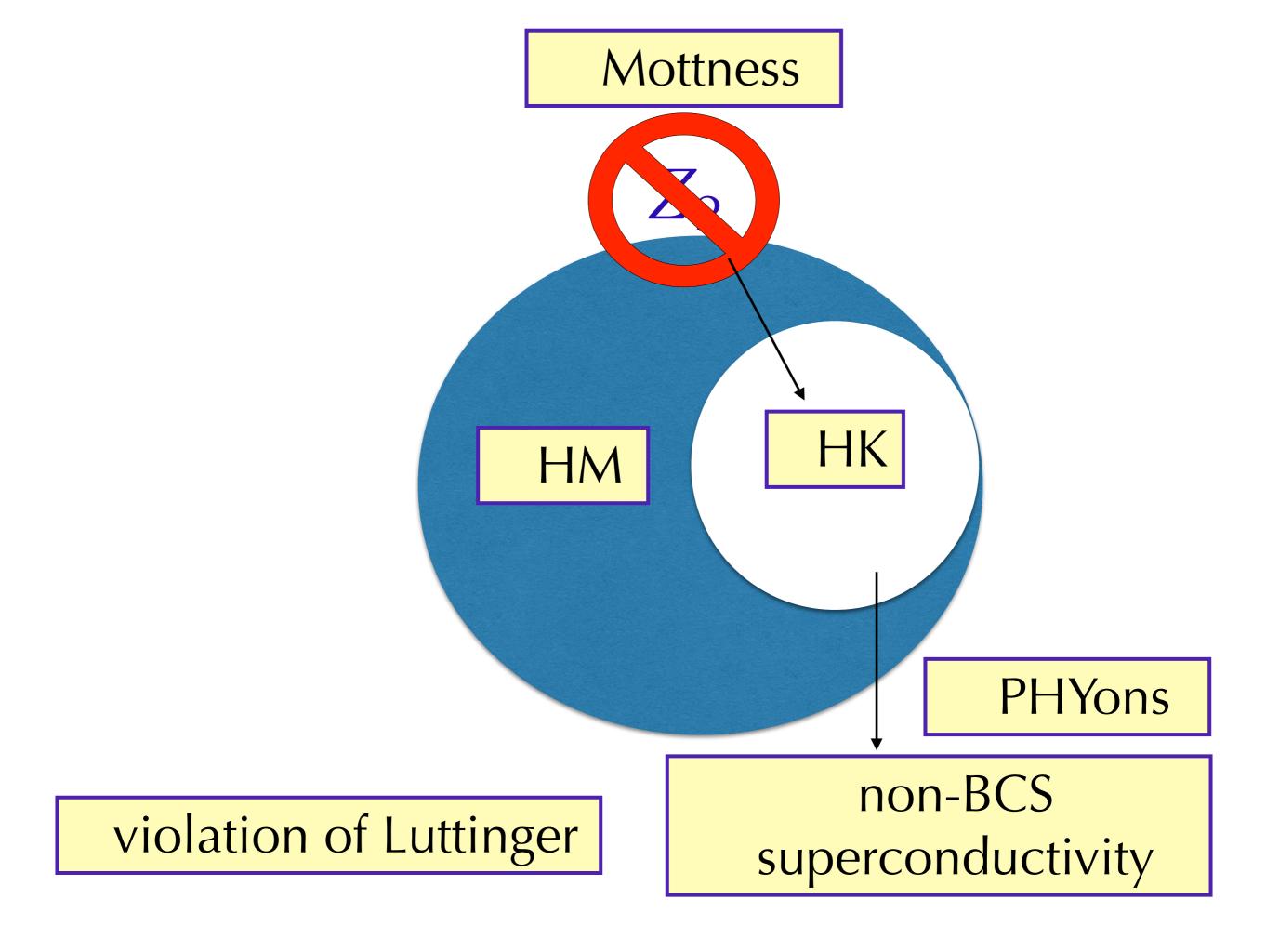
dynamical spectral weight transfer

is this the general mechanism of the color change?

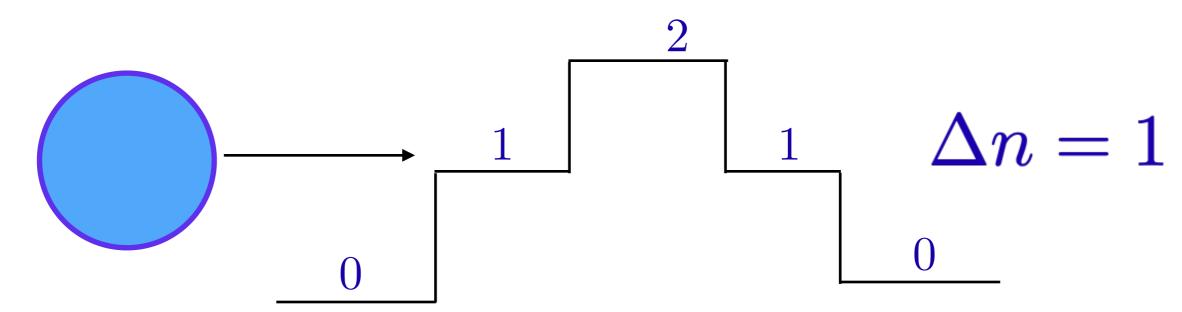
Superfluid Density

Mottness-induced suppression





HK and EFL?



other hand, we have

tween two regions with different $n(\mathbf{k})$ carries a Fermi surface with anomaly coefficient m given by the difference Δn . In a Fermi gas, $n(\mathbf{k})$ has a natural interpretation: it

$$\frac{2mq}{(2\pi)^2} \mathcal{V}_F = \rho \pmod{2}.$$
 (64) $\Delta n = m = 2$

The extra factor of two takes into account the two possible spin values, and agrees (setting m=1) with the usual result for spinful Fermi liquids. (Thus, if we had defined m with respect to the total charge as we did in Section VI, we would have found m=2 for a spinful Fermi liquid/EFL).

